

Energy Transfer by Inertial waves in Rotating Turbulence

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The context: Want to understand turbulence in **3D rotating systems**.
Atmosphere, Oceans, Flows within the Earth's mantle other planetary flows.

Students involved:
Itamar Kolvin
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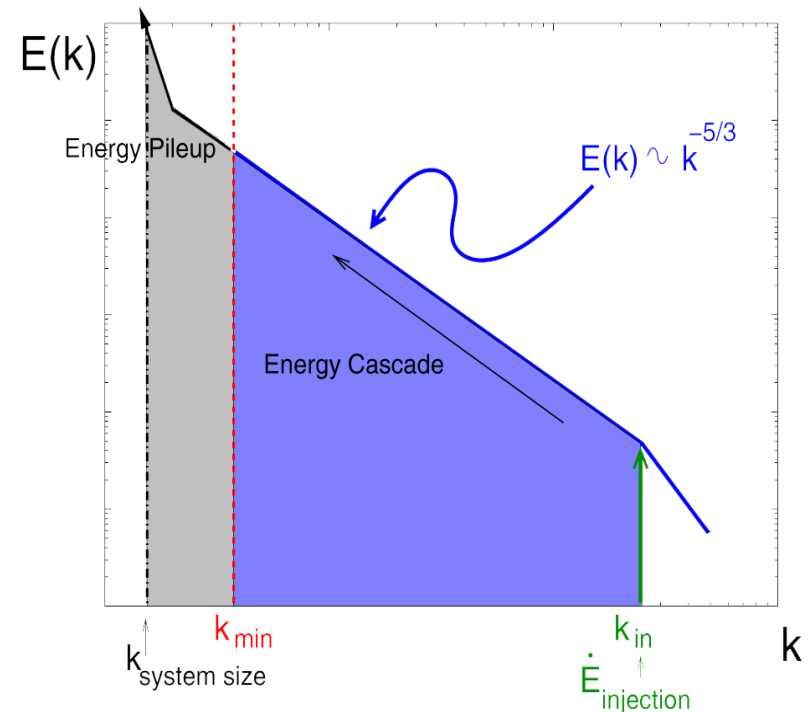
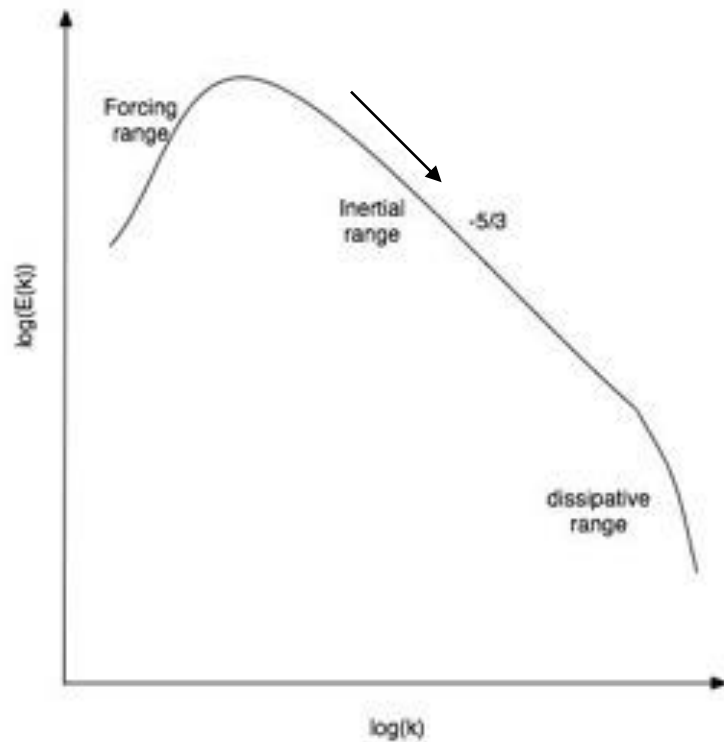
Outline

- The connection between rotating and 2D turbulence
- Inertial waves in rotating fluid
- Energy transfer in rotating turbulence
- Statistics
- A call for wave turbulence description

What is the right framework to describe 3D rotating turbulence?

3D isotropic turbulence – Richardson, Kolmogorov, forward energy cascade...

2D isotropic turbulence – Batchelor, Kraichnan, inverse energy cascade...



Rotating 3D turbulence is often described in terms of **2D turbulence**

A central question:

The equivalence and differences between rotating and 2D turbulence

The suggestion for equivalence is sometimes justified using **Taylor-Proudman's theorem**

$$\cancel{\frac{\partial \vec{u}}{\partial t}} + \vec{u} \cancel{\cdot} \nabla \vec{u} = -\frac{\nabla p}{\rho} + \cancel{v} \nabla^2 \vec{u} - 2\vec{\Omega} \times \vec{u} \quad \Rightarrow \quad \vec{\Omega} \cdot \nabla \vec{u} = 0$$

For $\vec{\Omega} = \Omega_z$ We have: $\frac{\partial}{\partial z} \vec{u} = 0$ **Quasi 2D**

Another question:

What is the mechanism that maintains two dimensionality of the flow?
(How energy and momentum are transferred along the axis of rotation?)

$$\frac{\partial \vec{u}}{\partial t} + \vec{u} \times \nabla \vec{u} = -\frac{\nabla p}{\rho} + \nu \nabla^2 \vec{u} - 2\vec{\Omega} \times \vec{u}$$

For small and slow perturbations to fluid at rest (in the rotating system)

Coriolis driven inertial waves propagate along the axis of rotation (Greenspan 1968) with group velocity:

$$\mathbf{v}_g = 2\bar{\mathbf{k}} \times (\bar{\boldsymbol{\Omega}} \times \bar{\mathbf{k}}) / |\bar{\mathbf{k}}|^3$$



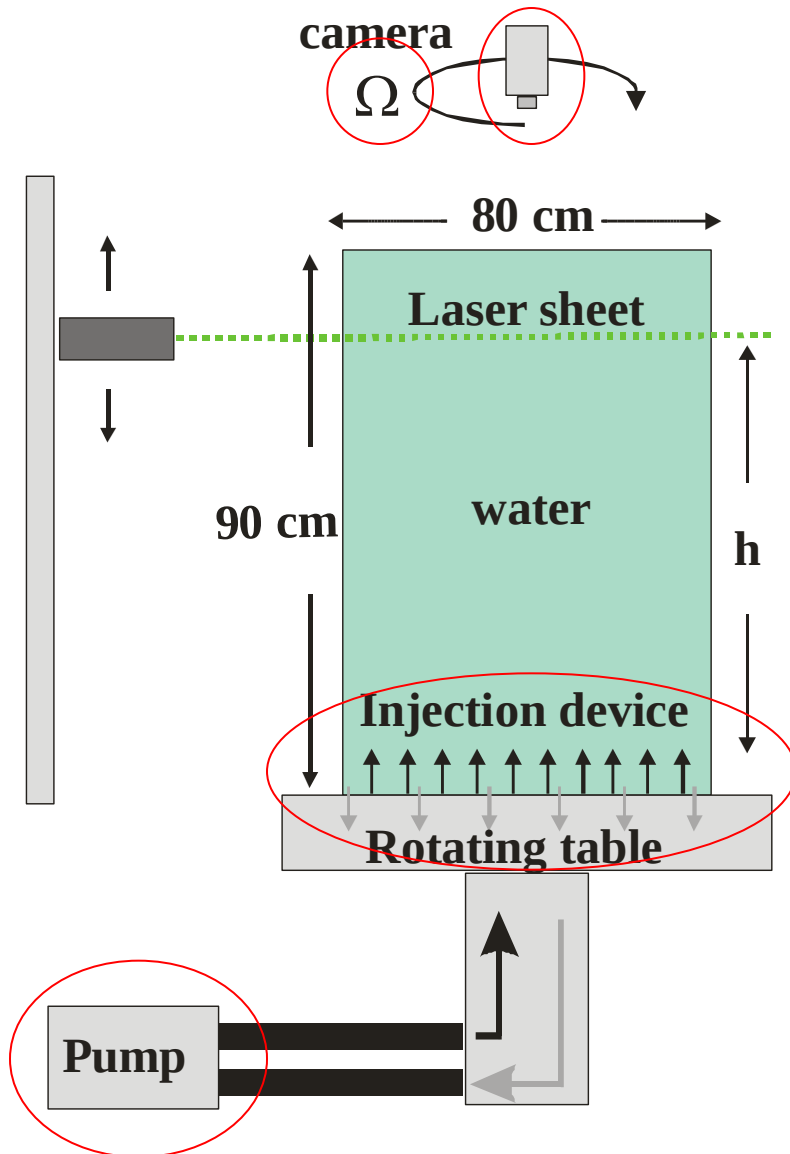
$$|\mathbf{v}_g(k)| = \frac{2\Omega}{|k|}$$

The frequency of inertial waves must satisfy:

$$\omega \leq 2\Omega$$

Are these waves important in turbulent rotating flow?

Experimental system



Ω up to 16 Rad/s

Max. flow rate 3 L/s $\Rightarrow$ $\sim$ 300 W

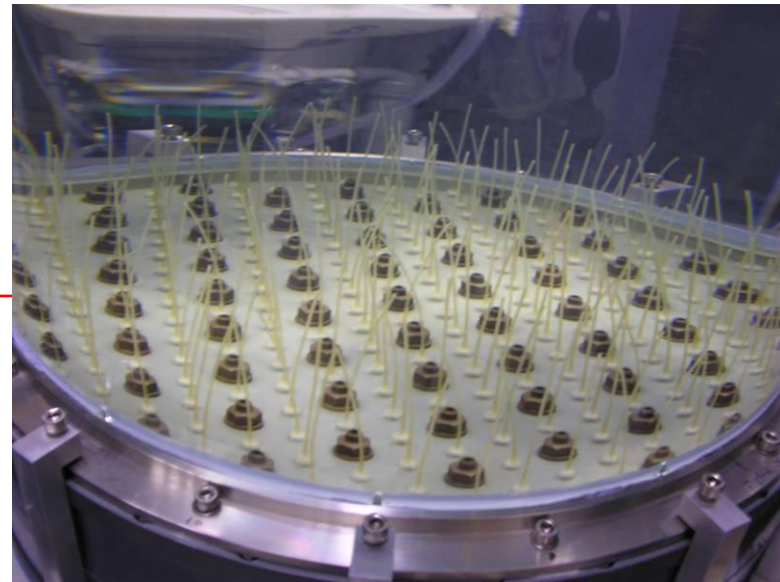
250 outlets and 70 inlets in
hexagonal lattice

1 Mpix, 30 f/s, with less than
1ms "dead time", for PIV
measurements

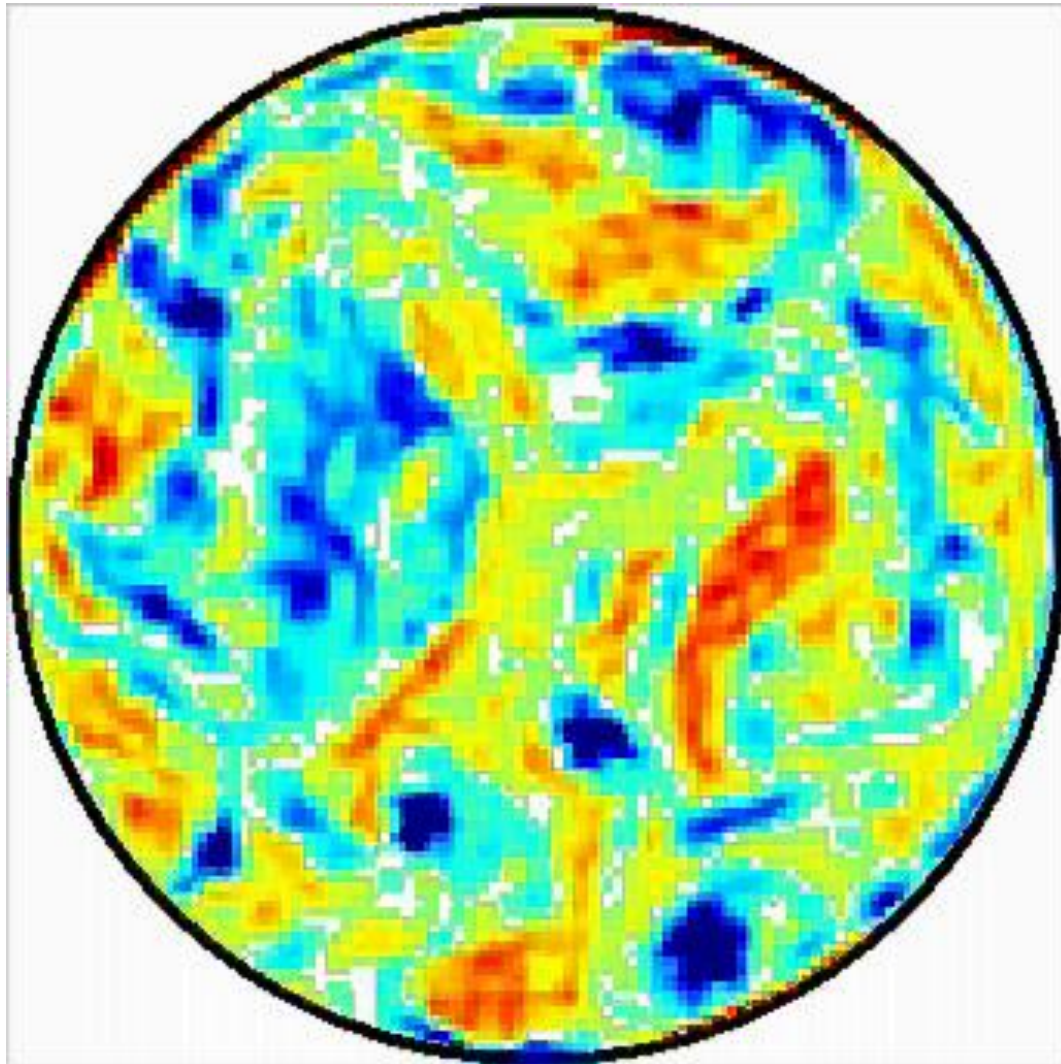


Camera

Injection Nozzles



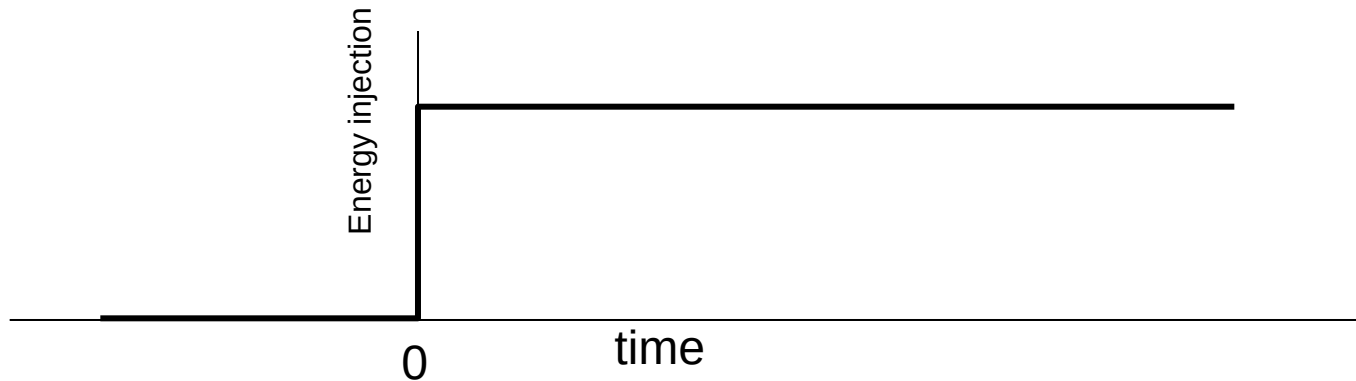
In a steady state (vorticity field)



Experiment 1

The system is brought to a solid body rotation ($u=0$) at a given rotation rate Ω .

At $t=0$, we start injecting energy at a given flow rate (generating a **step function in the injected power**)

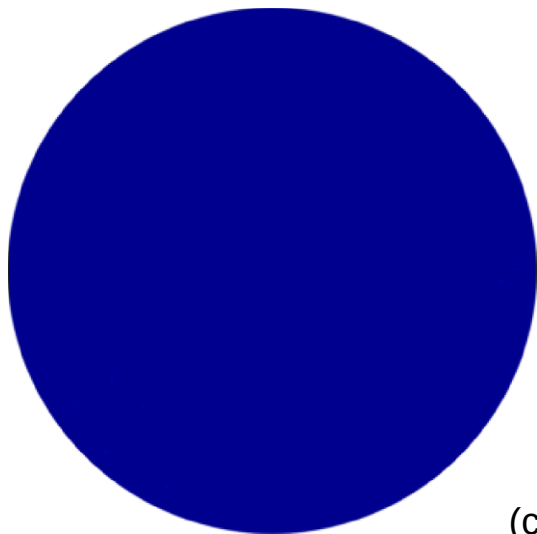


We measure the horizontal velocity (u,v) field **at height H**

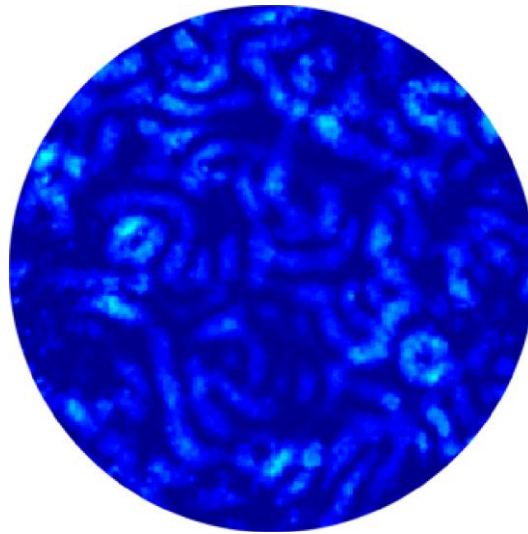
Deriving energy power spectrum, $E(k)$ and energy density “map”, (u^2+v^2) , as functions of time.

Energy Density evolution

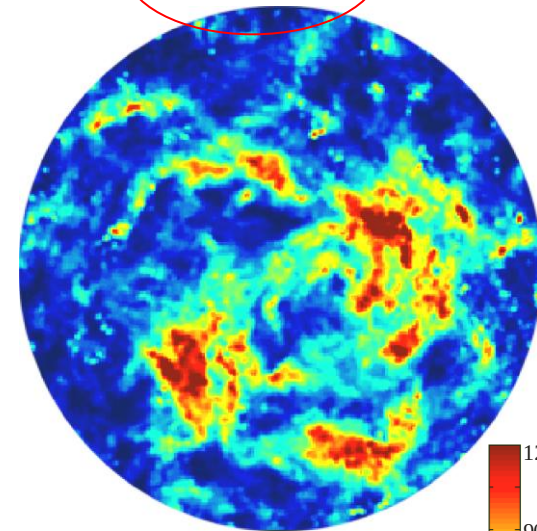
t=1.3 s



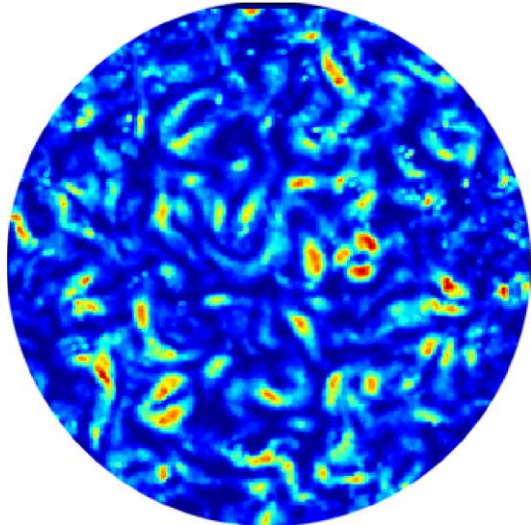
t=3.3 s



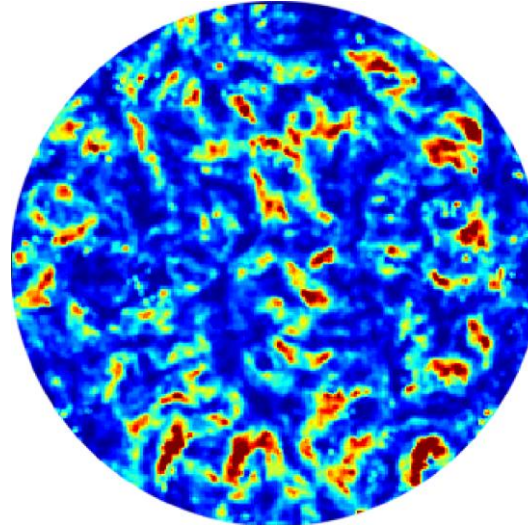
t=100 s



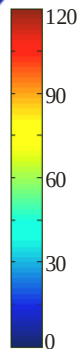
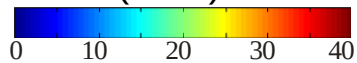
t=4.7 s



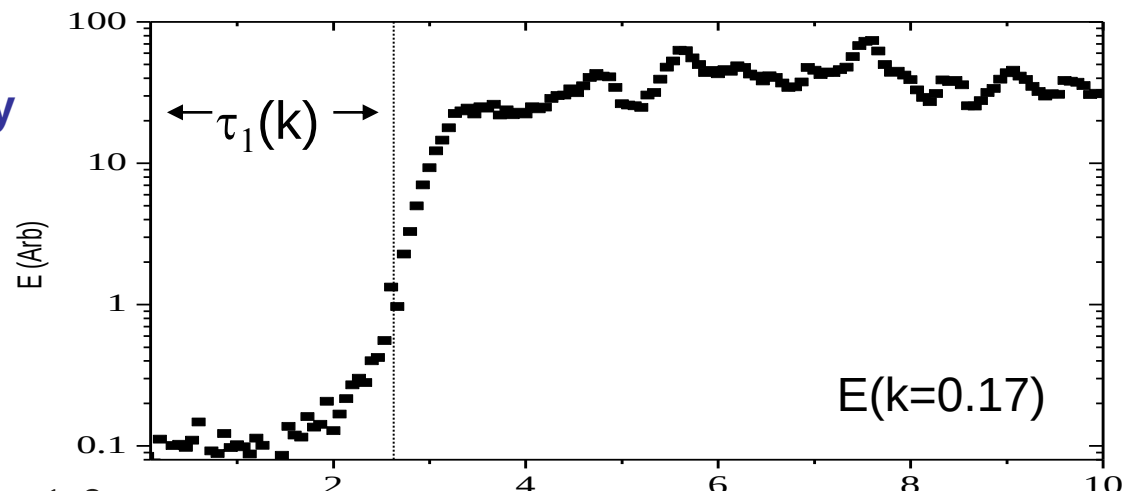
t=10 s



(cm/s)²



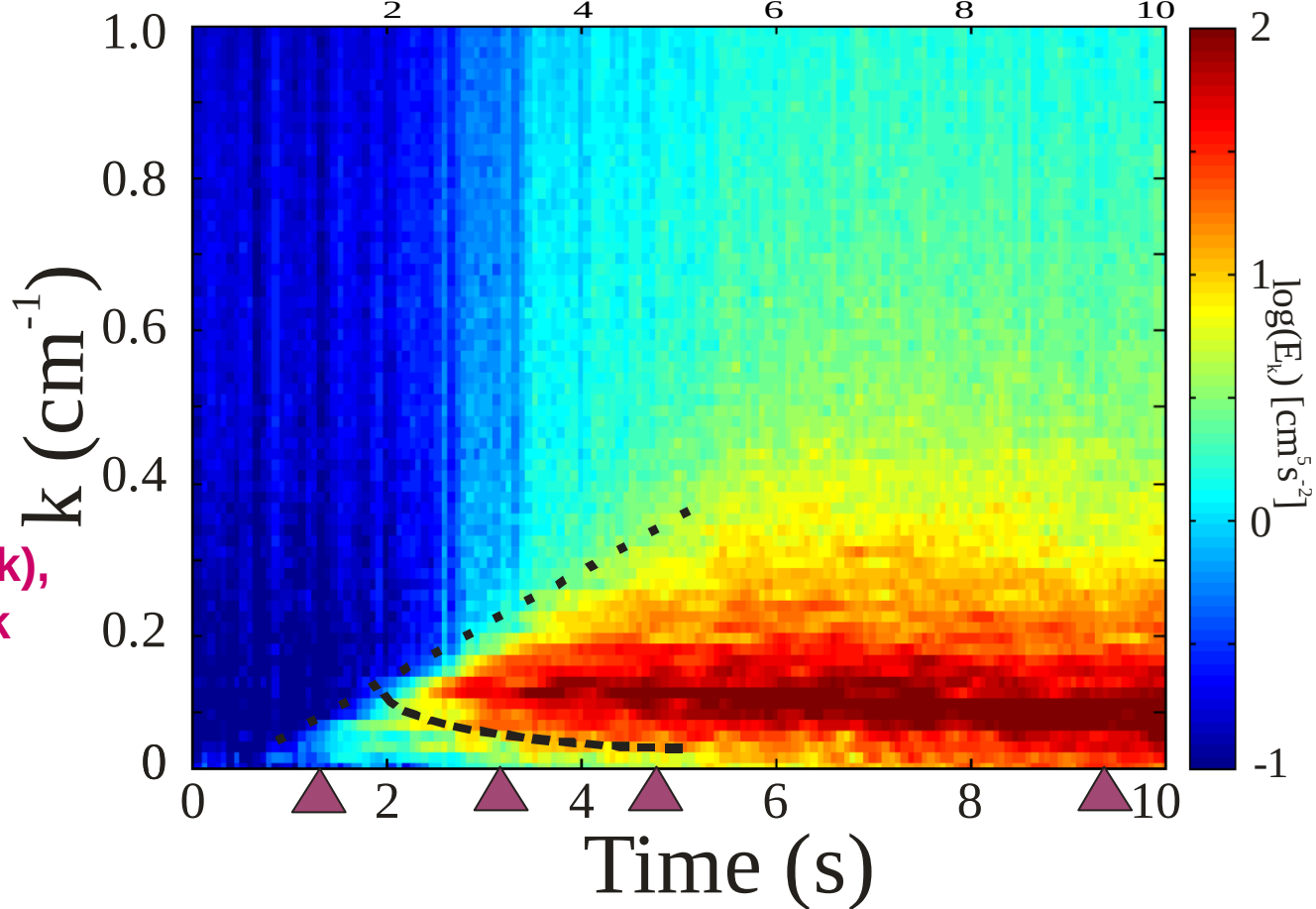
Sharp “arrival” of energy



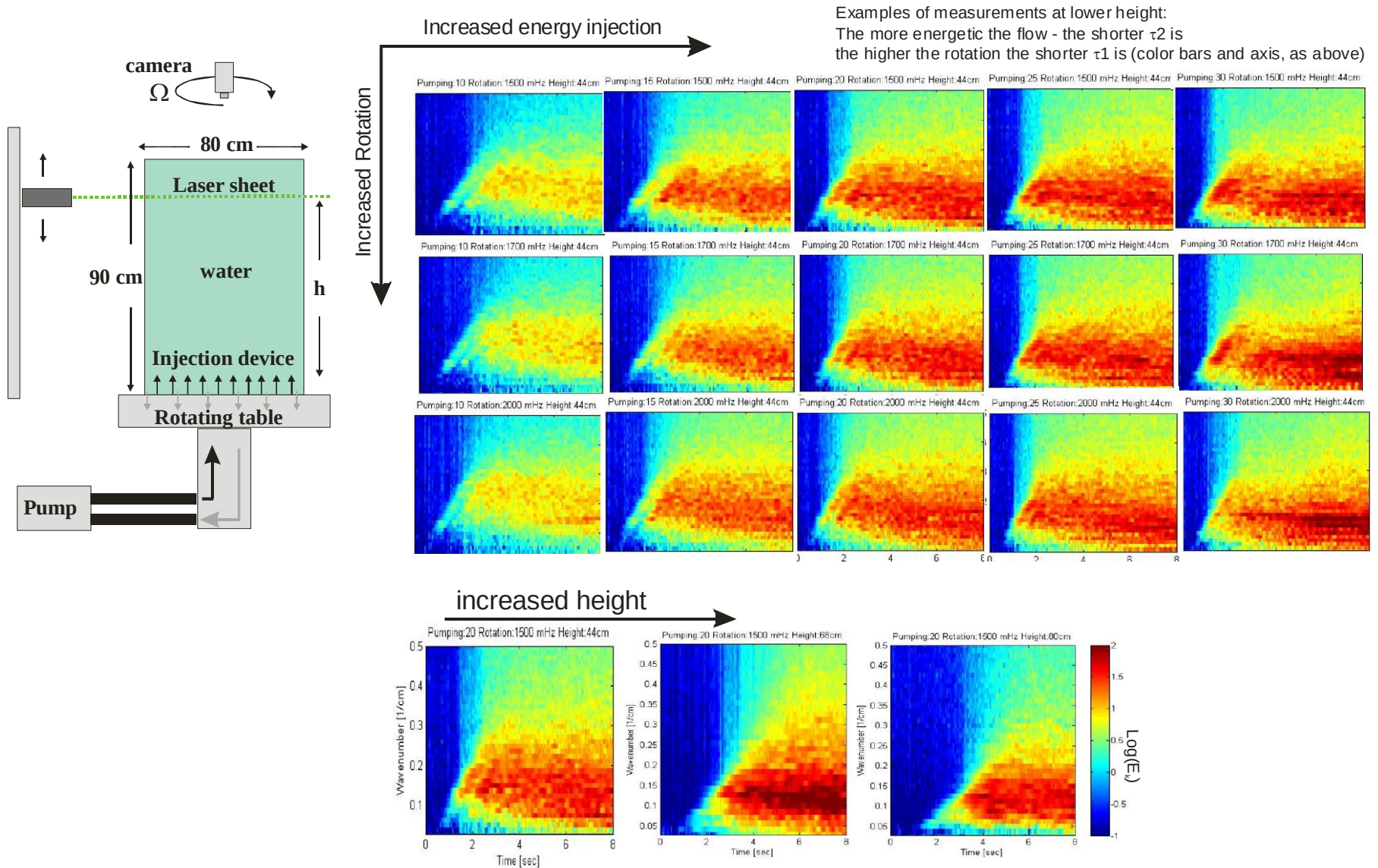
Two “Populating fronts”
in the spectrum

One is linear in k ,
defining $\tau_1(k)$

The second defines $\tau_2(k)$,
which decreases with k

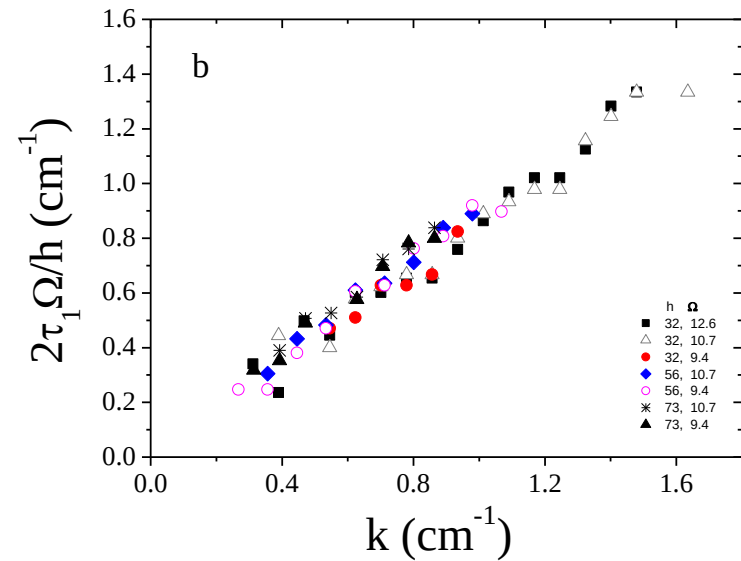
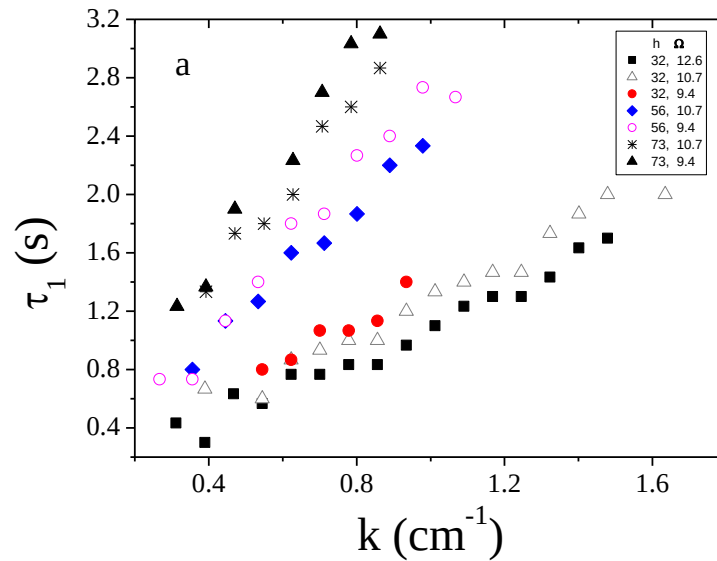


Variation of fronts properties with rotation, energy injection and height



Scaling of τ_1

The energy carried by an **inertial wave** of wave number **k** would arrive at **h** by a time: $t = \frac{h}{v_g} = \frac{hk}{2\Omega}$



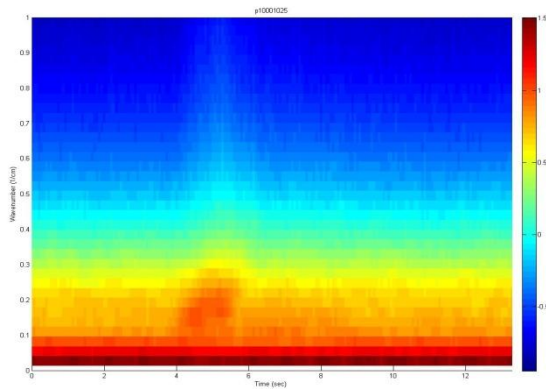
τ_1 is the **traveling time** of inertial waves to the measuring plane



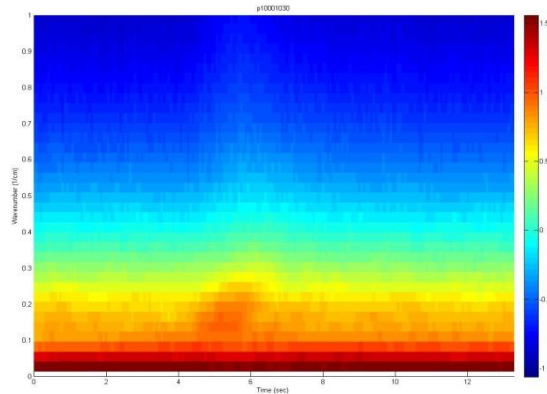
All the energy transfer along z is done by inertial waves

Waves propagate within existing turbulence

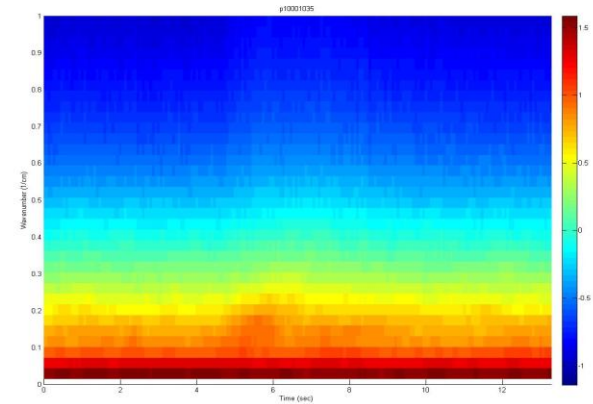
H=25 cm



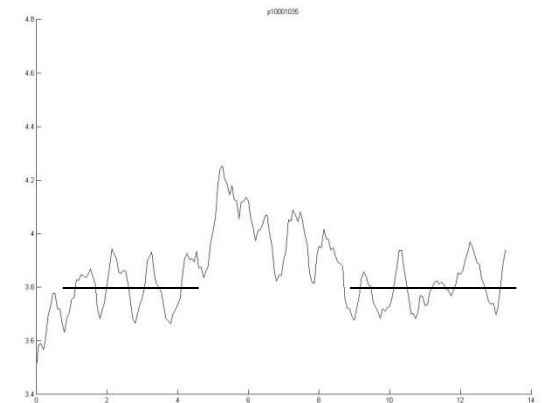
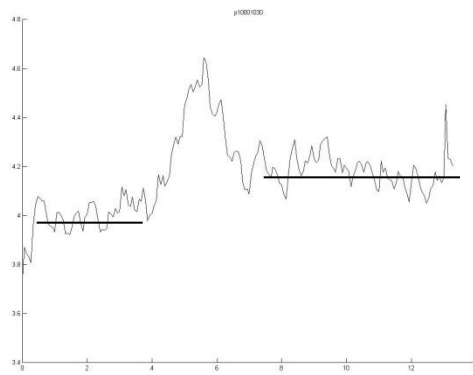
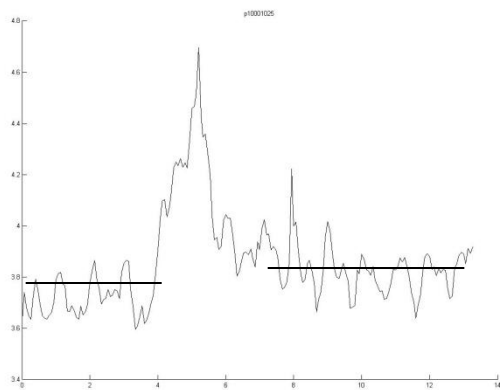
H=30 cm



H=35 cm



Energy spectrum vs. time

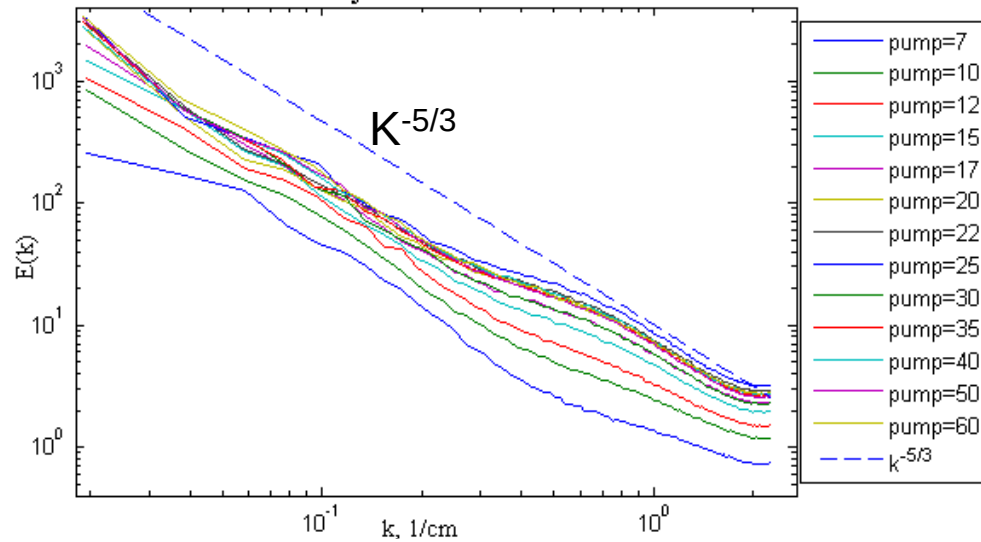


Total Energy vs. time

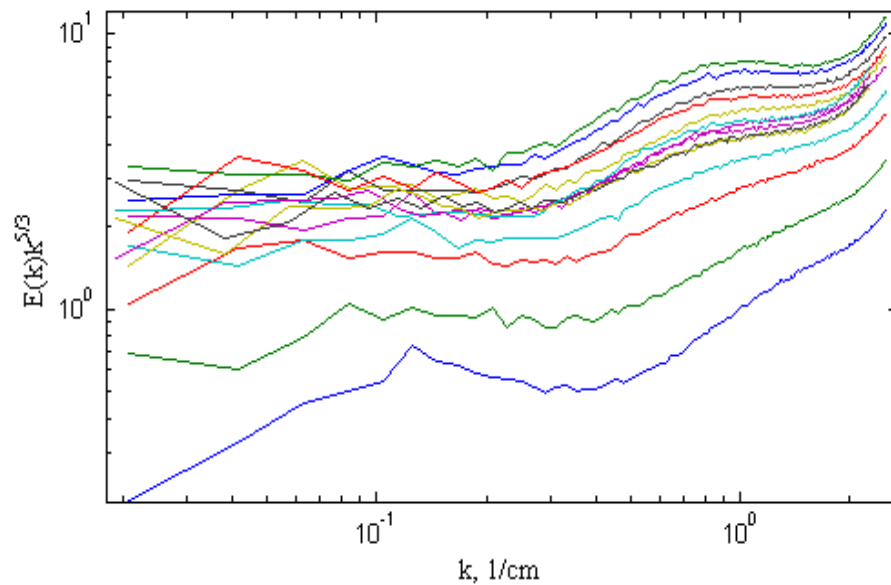
Are inertial waves significant in steady state?

Energy spectrum seems consistent with 2D turbulence

Steady state for $\Omega = 1.5$ Hz



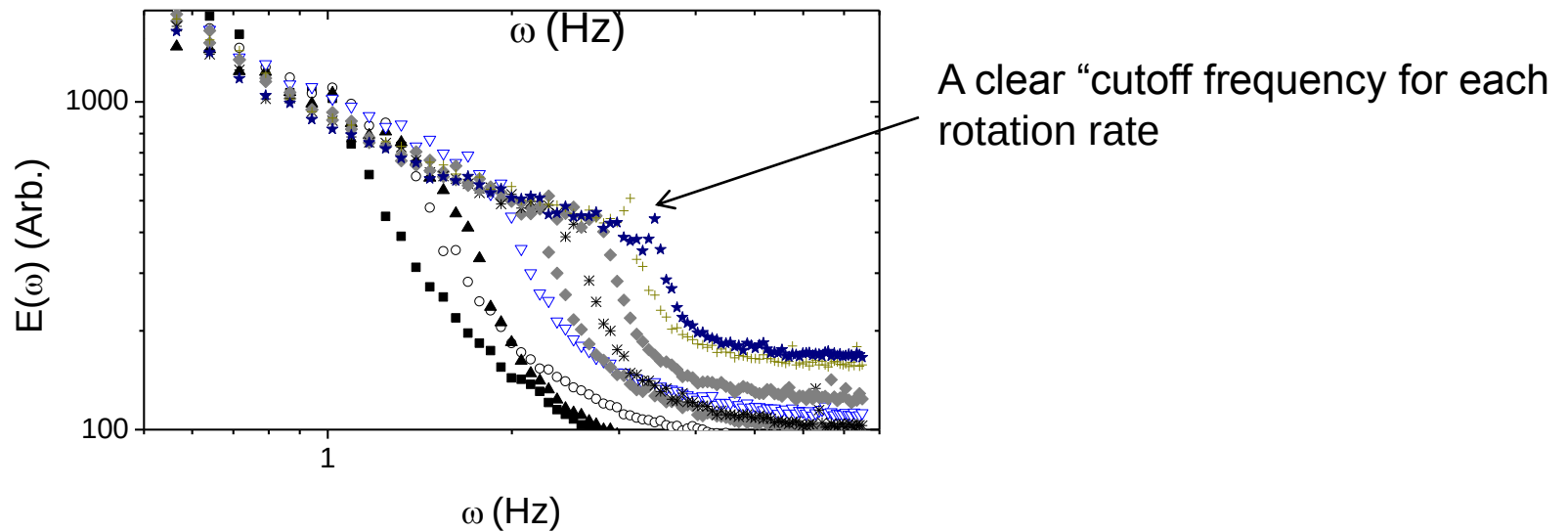
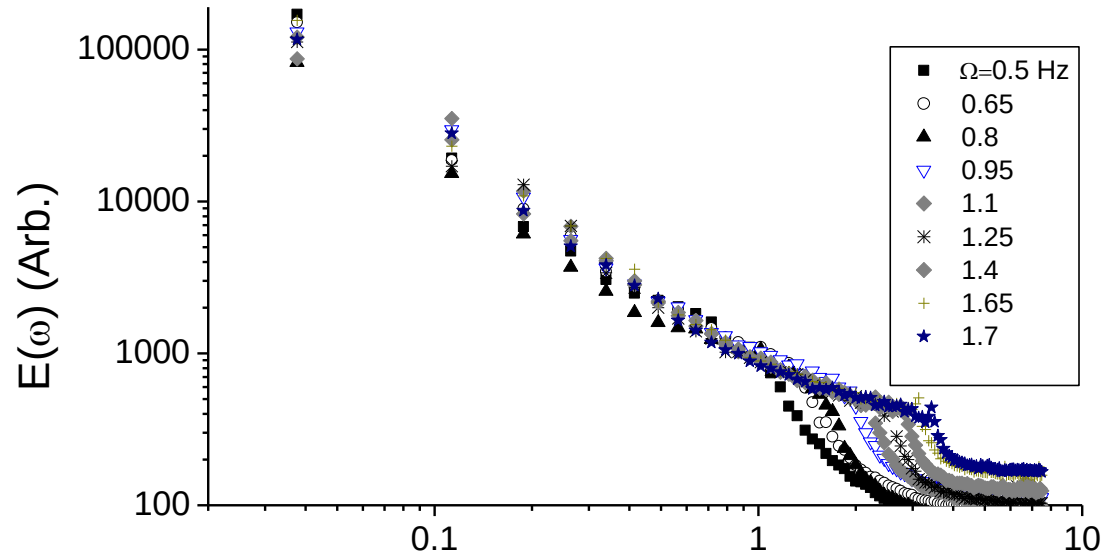
Steady state for $\Omega = 1$ Hz



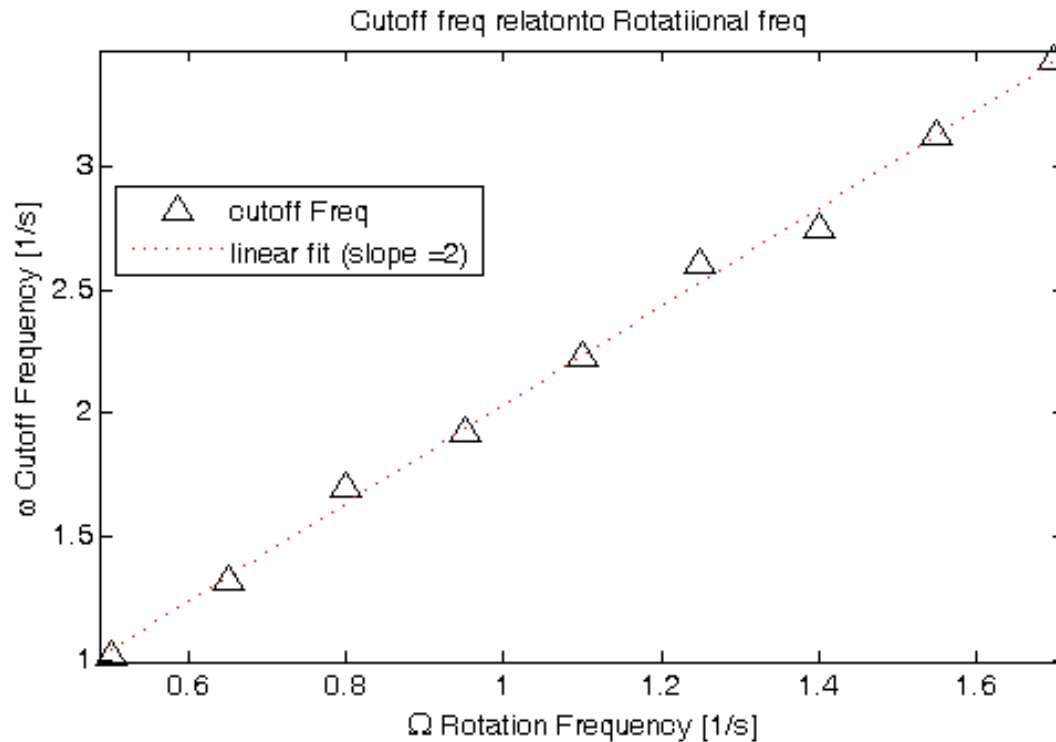
“compensated” spectrum

Measurements by Yuval Vardi

The **temporal** energy spectrum for different rotation rates



The cutoff frequency vs. Ω



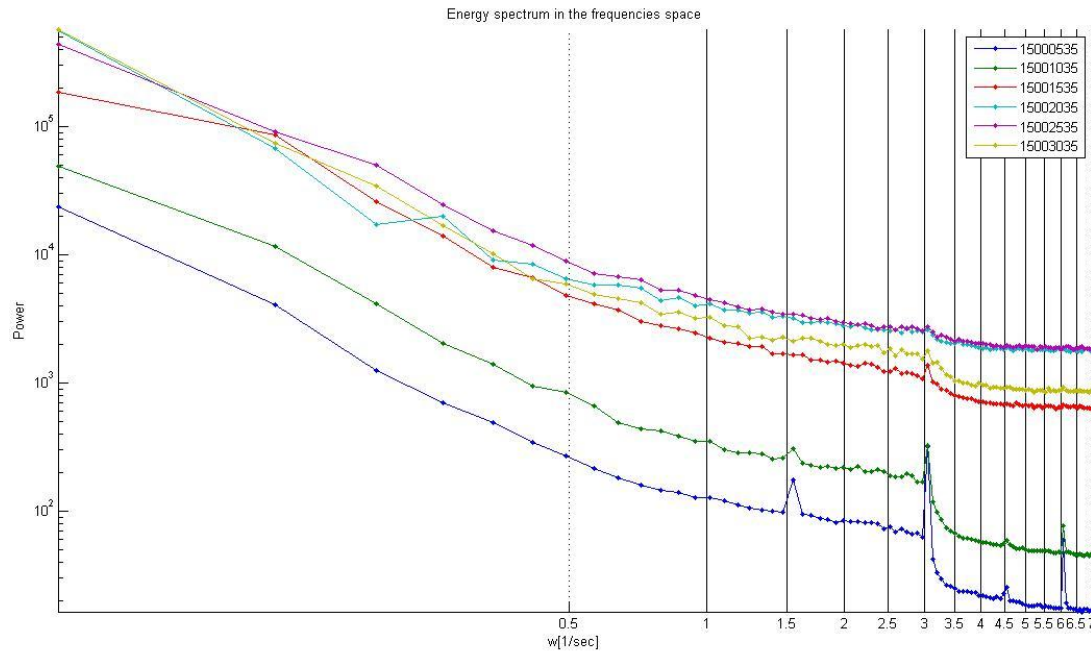
$$\omega_{\text{cutoff}} = 2\Omega$$

Consistent with the cutoff in the dispersion relation of inertial waves:

$$\omega \leq 2\Omega$$

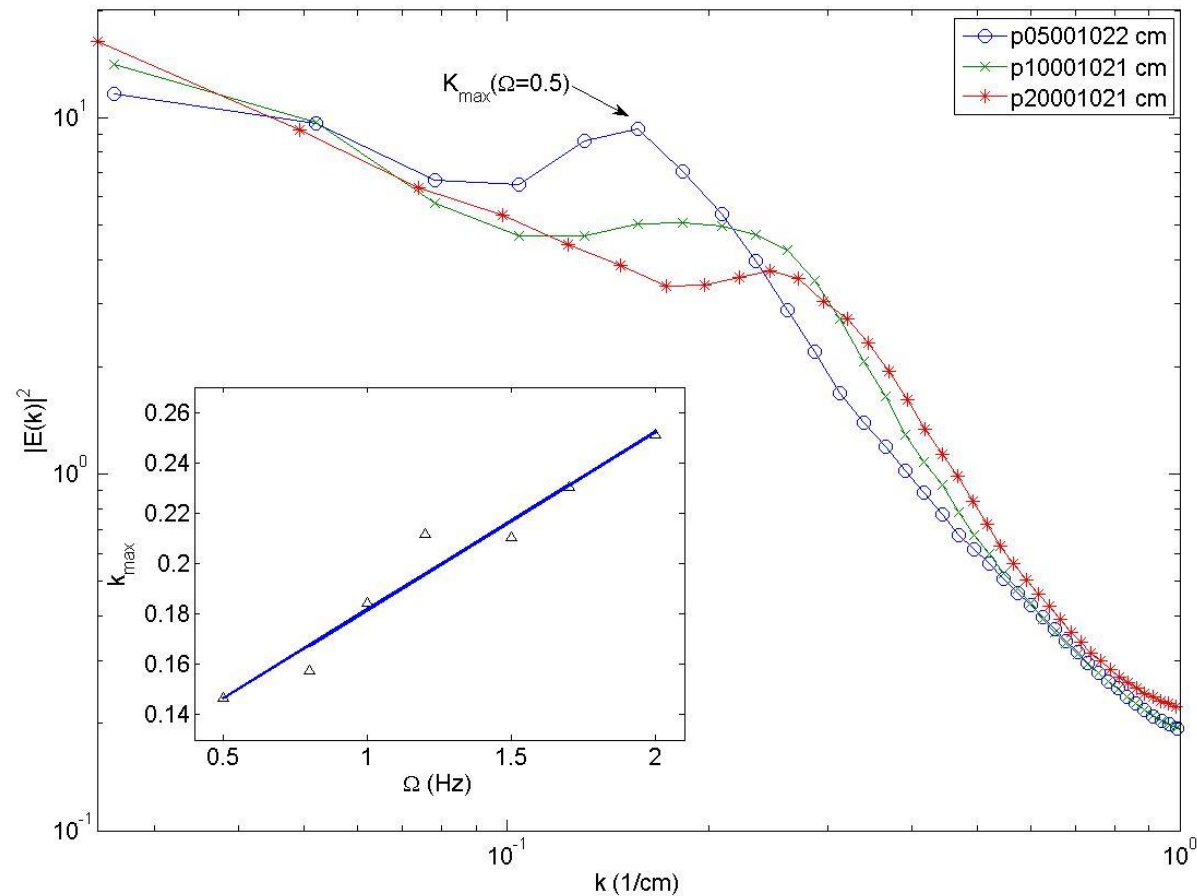
Suggests that inertial waves strongly affect the statistics even in a steady state

The **temporal** energy spectrum for different pumping rates



What is the dependence on the dissipation rate?

The injected energy spectrum at different rotations (same pumping rate)



A shift of $k_{\max}$ with Ω

Conclusions

High amplitude inertial wave packets can propagate in a turbulent rotating fluid.

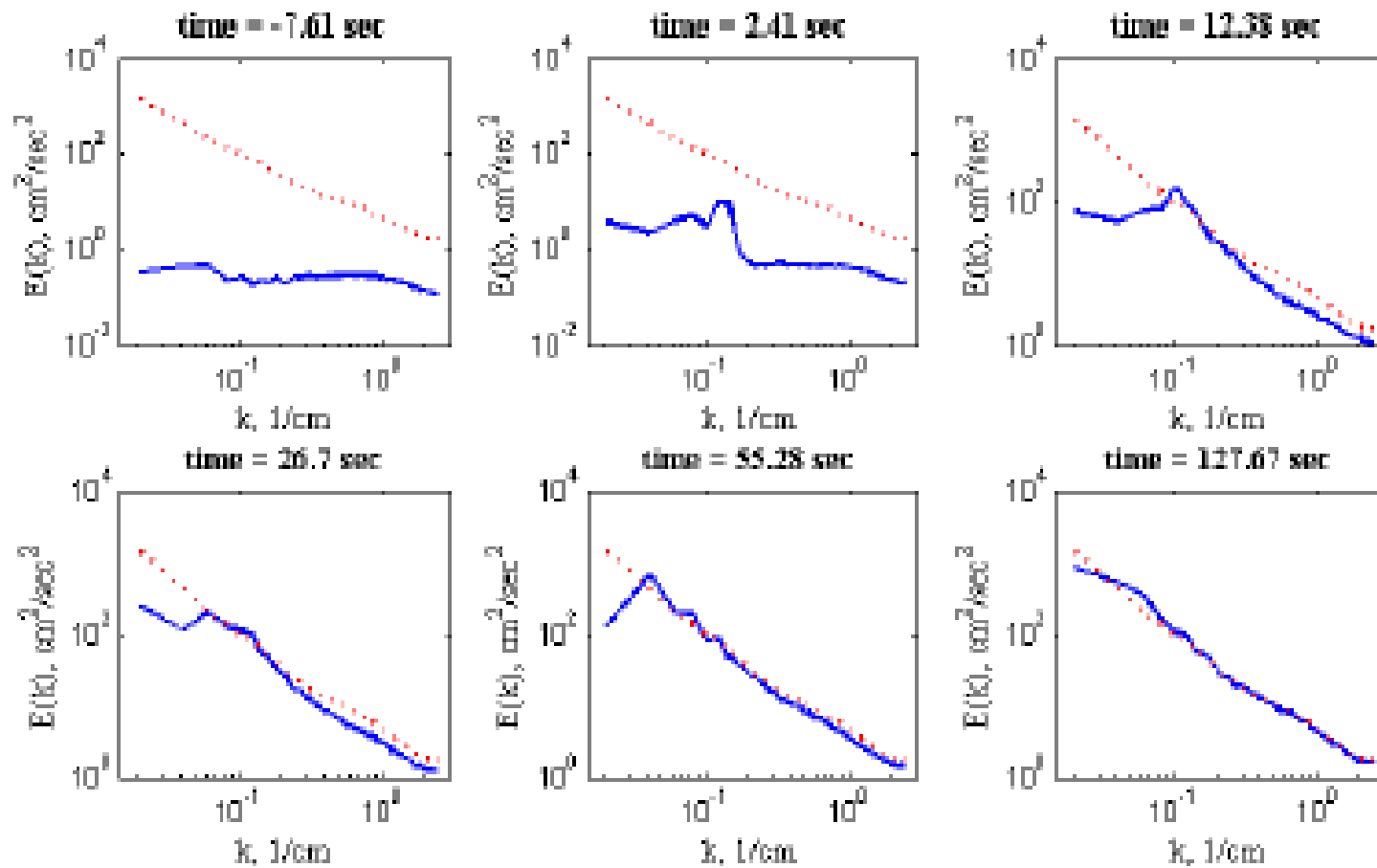
They propagate with the group velocity calculated for small perturbations

In the flow that was studied, **energy is transported by inertial waves**

Statistics of rotating turbulence, contains “finger prints” of inertial waves.

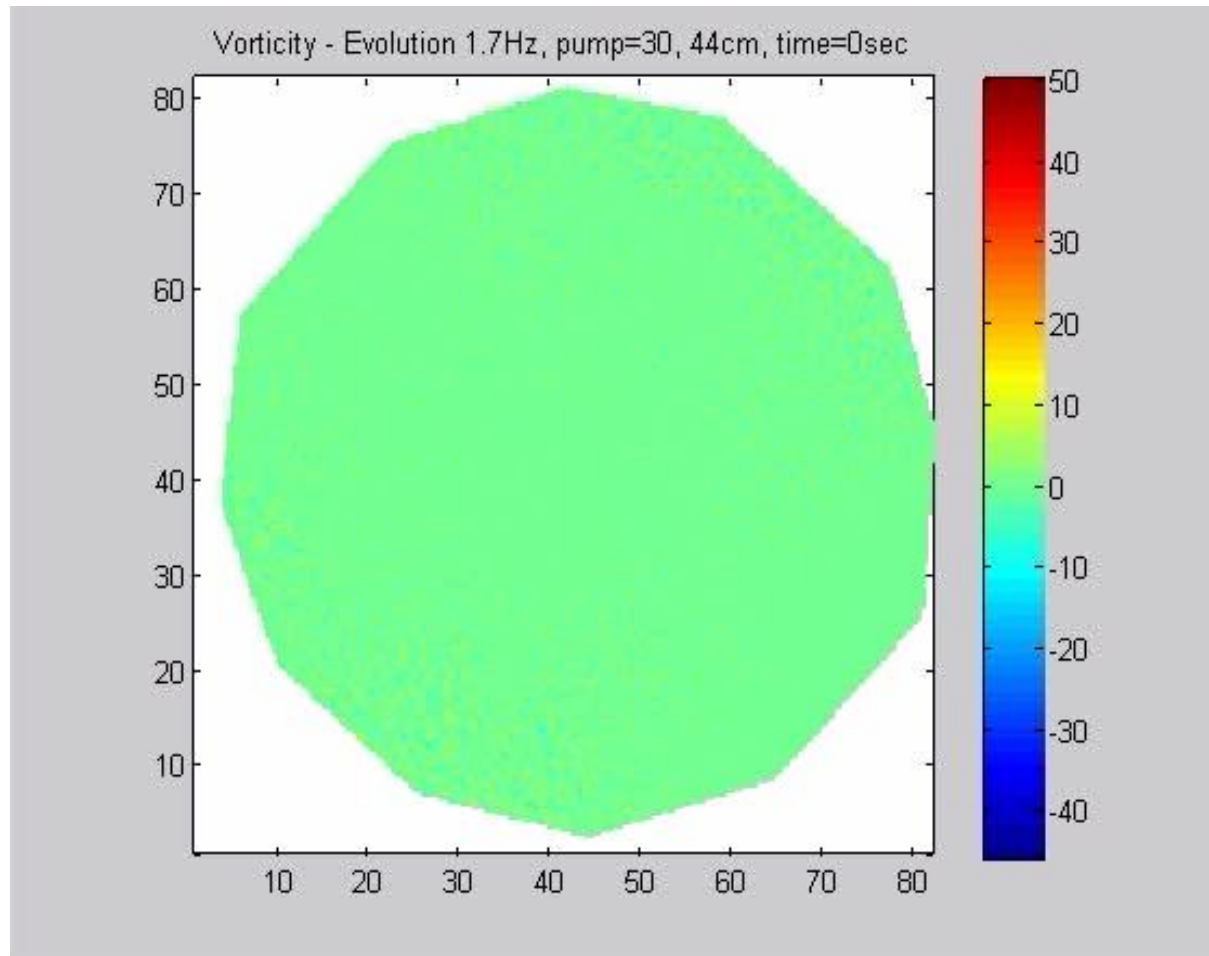
Can we write an “**inertial wave turbulence**” description?

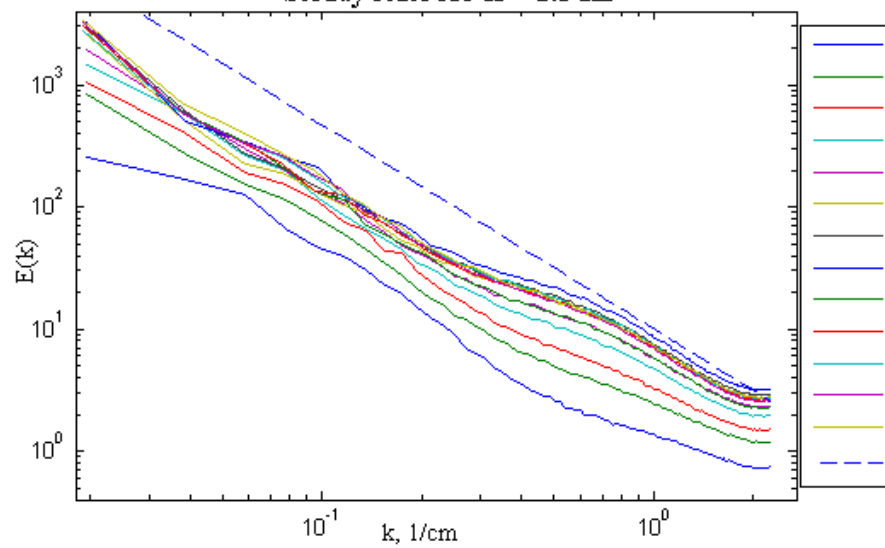
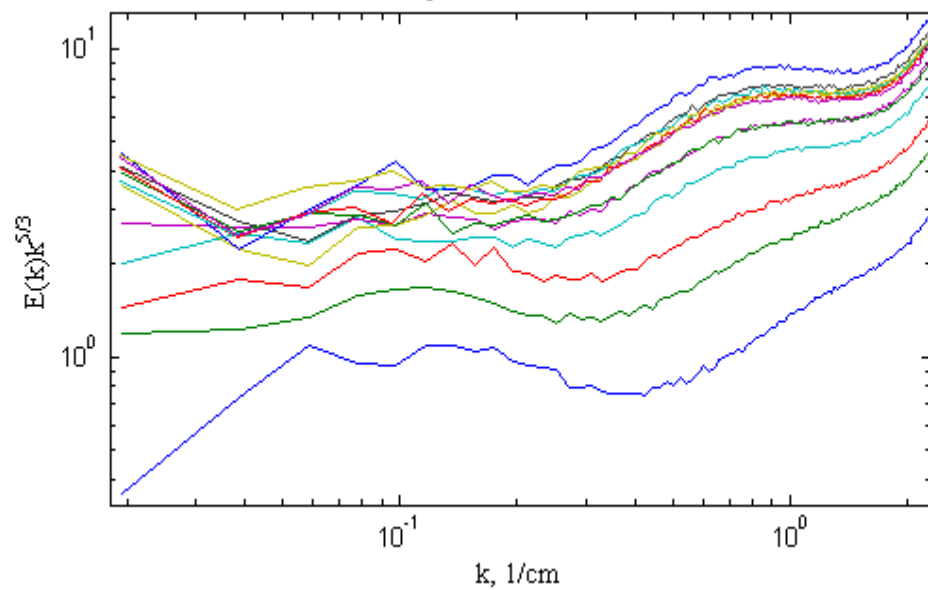
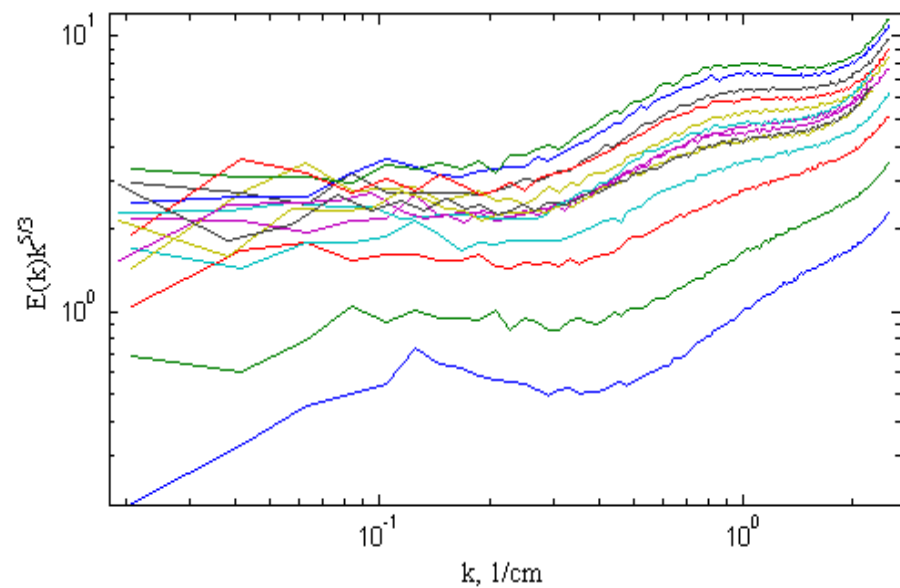
Spectrum evolution over longer times



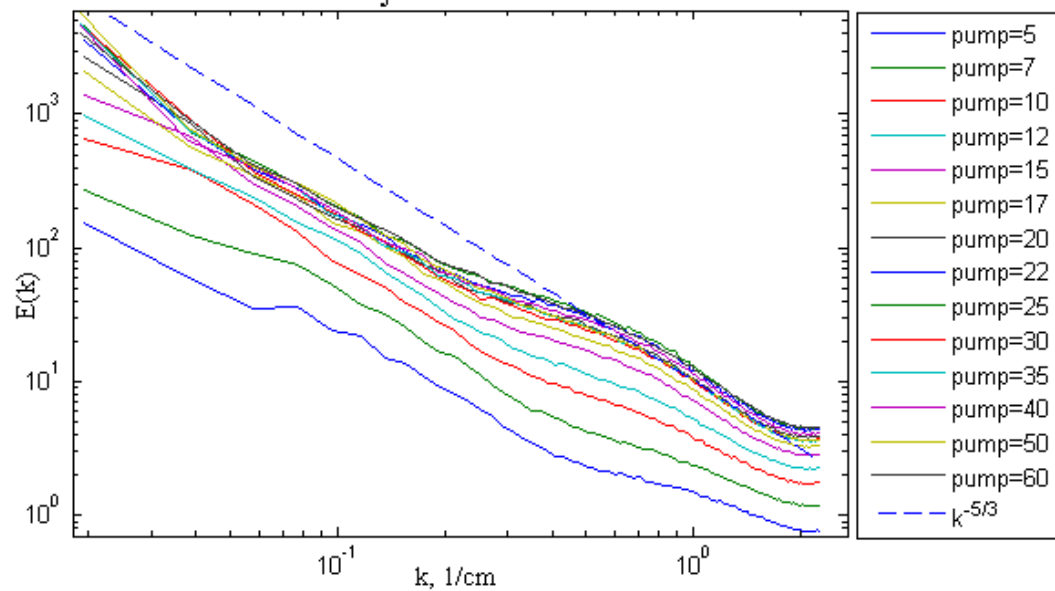
Vorticity

$$\omega = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}$$

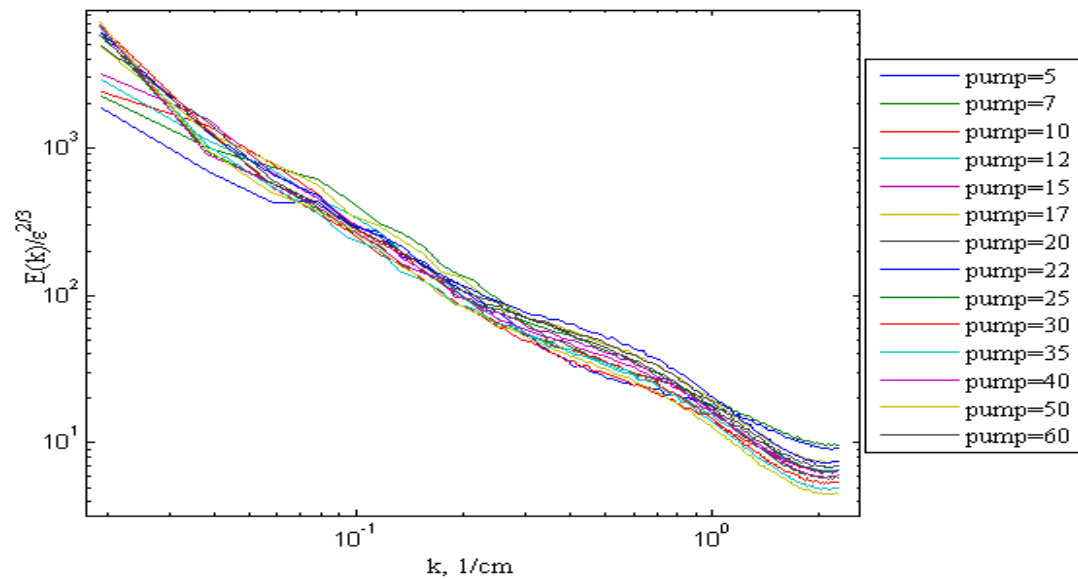


Steady state for $\Omega = 1.5$ HzSteady state for $\Omega = 1.5$ HzSteady state for $\Omega = 1$ Hz

Steady state for $\Omega = 2$ Hz



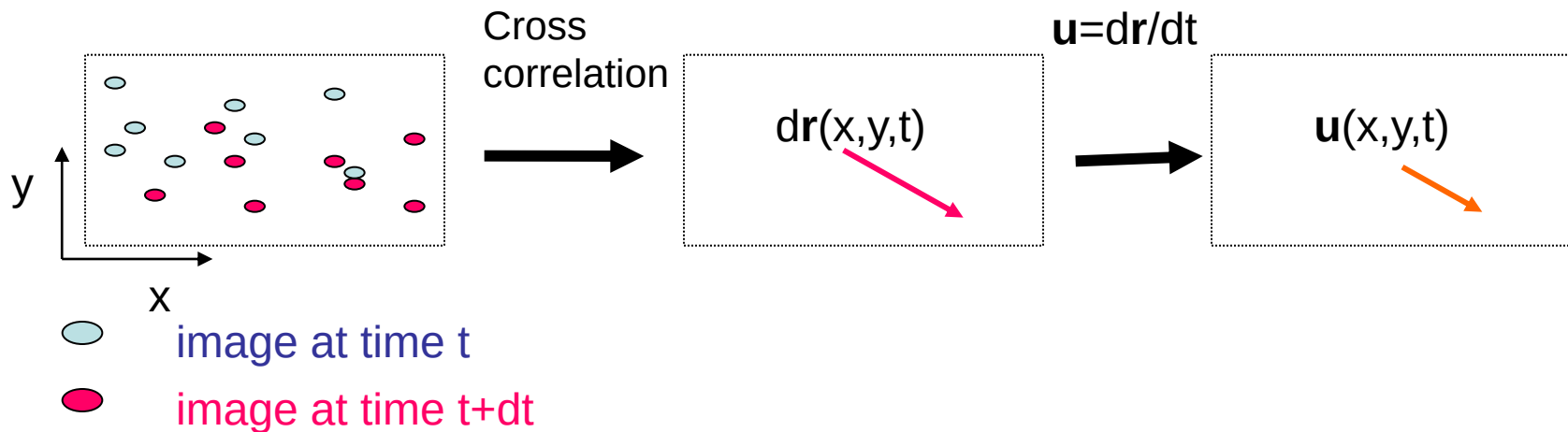
Steady state collapse for $\Omega = 2$ Hz



Scaling by $\epsilon^{2/3}$

PIV system

- Using $\sim 50 \mu\text{m}$ Polystyrene particles ($\rho = 1.06 \text{ g/cm}^3$) for light scattering
- Dividing an image to $\sim 200 \times 200$ “cells”, each with ~ 10 particles
- Calculating the cross correlation between images of different times, t and $t+dt$, within each cell



- Image 1 Megapixel, 10 bit
- Shortest $dt \sim 1 \text{ ms}$ (by strobbing the laser)
- $u_{\text{Max}} \sim 1 \text{ m/s}$

Suggested scenario

$\tau_1 \sim Hk/\Omega$ – linear propagation time

$\tau_2 \sim (vk)^{-1}$ – nonlinearities of inertial waves

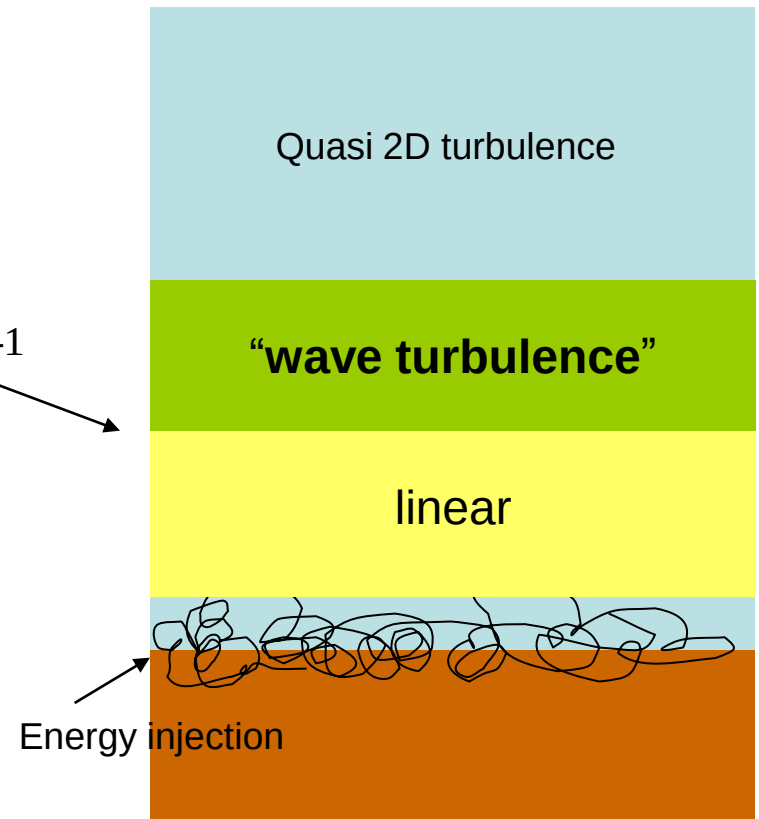
$\tau_3 \sim ?$ – time to achieve steady state

We have **dominance of linear behavior** for

$$k < \left(\frac{\Omega}{Hv} \right)^{1/2}$$

Or, for heights smaller than $H^* \sim \Omega k^{-2} v^{-1}$

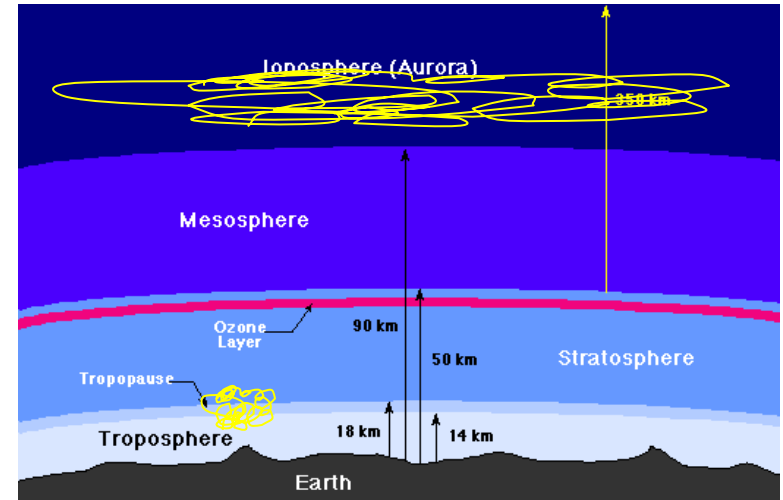
Above this height we expect to find a region dominated by (resonant?) nonlinear interaction of inertial waves – **“wave turbulence”**?



“Shooting” turbulence

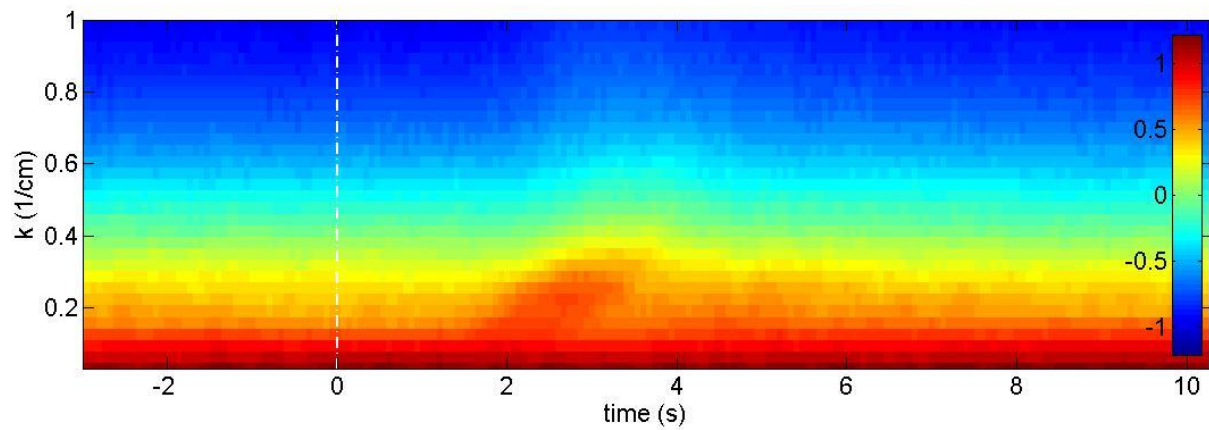
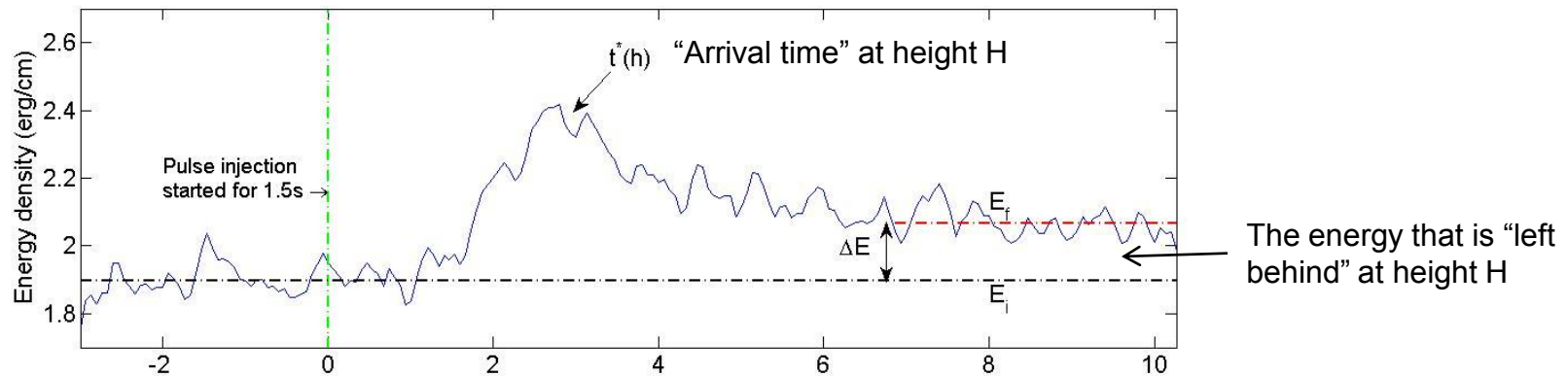
What happens in a turbulent atmosphere when there is a “sudden” blast of energy from a source?

Intuitively we would guess that most of the energy will accumulate **near the source**

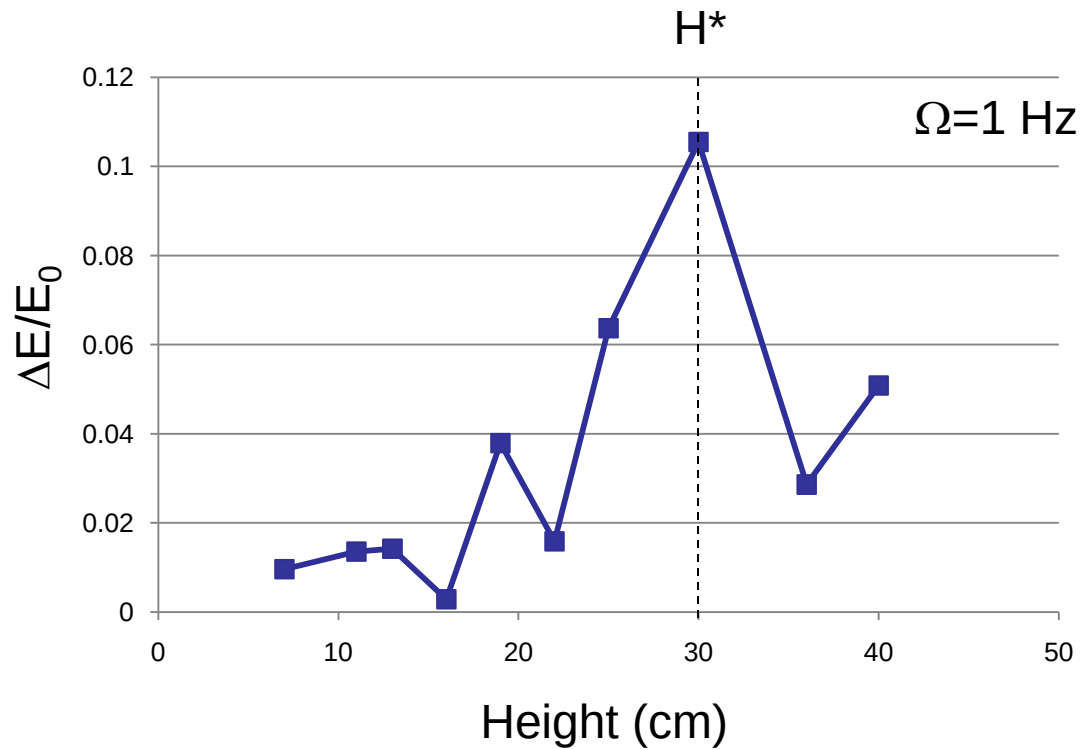


However, now we can predict, that the energy will **propagate linearly upwards** and will be **transferred to the turbulent field only at height larger than H^***

Measuring the propagating pulse



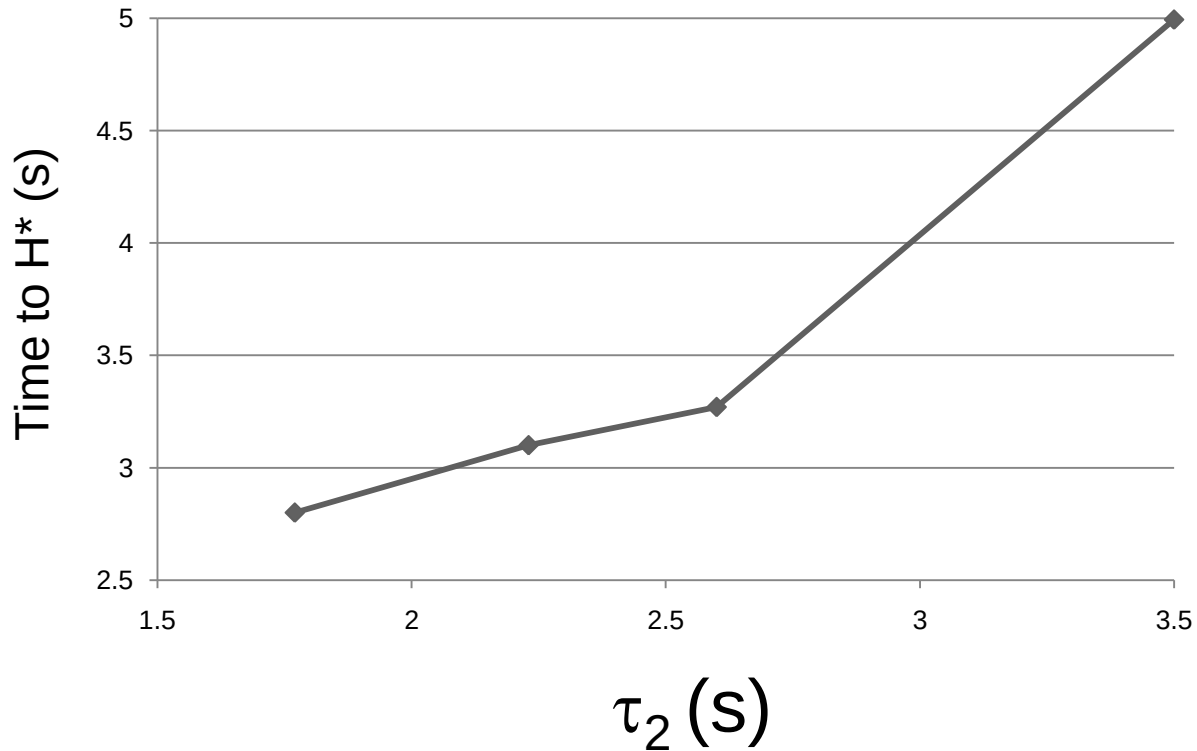
The energy “left behind” by the pulse vs. height



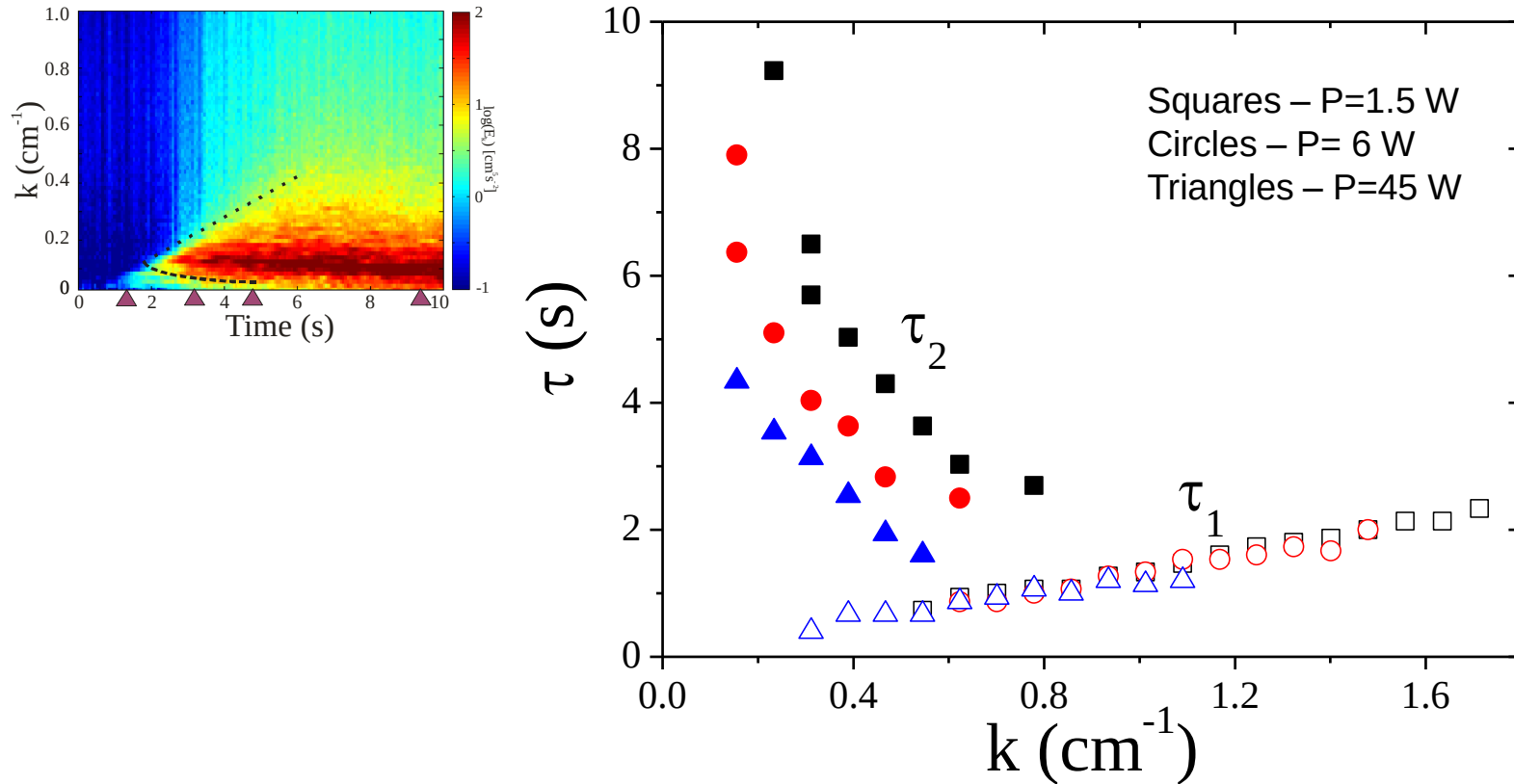
The energy of the pulse is concentrated at a selected height H^*

H^* is determined by τ_2

Find the traveling time to H^* under different conditions



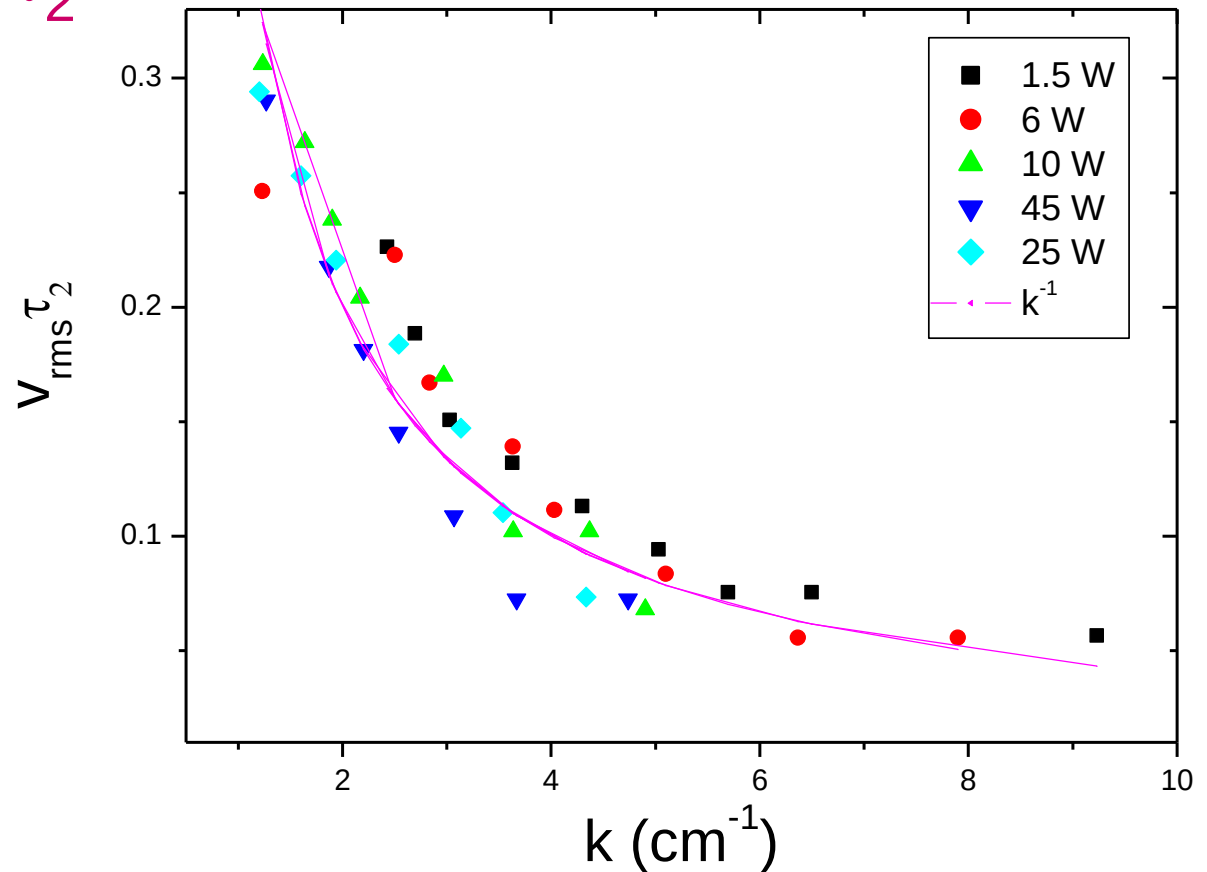
Dependence of τ_1 and τ_2 on injected Power - P



τ_1 is independent of P (as expected from a linear mechanism)

τ_2 decreases with P (An indication of the nonlinear nature of the underlying process)

Scaling of τ_2

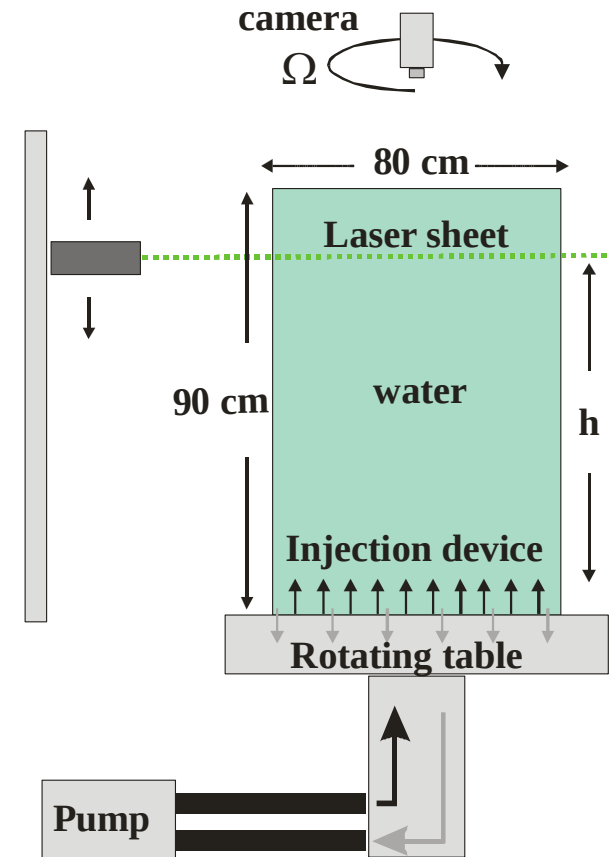
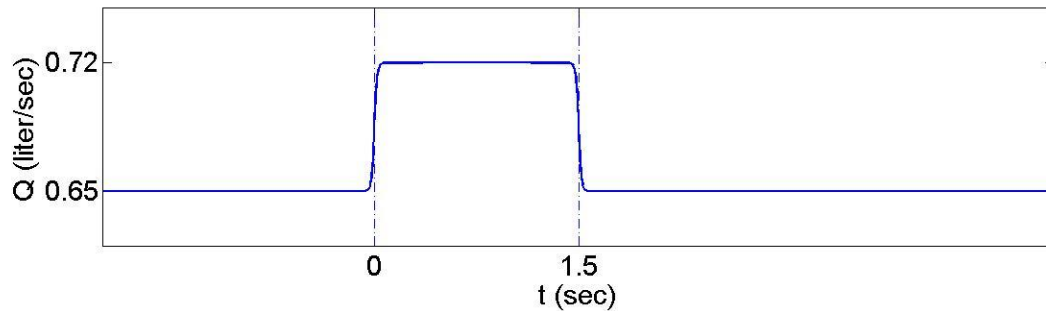


τ_2 scales roughly as the “eddy turnover time” - $(\nu k)^{-1}$

Should remember that steady state is achieved at times much longer than τ_2 — is there a third time scale?

The experiment

Injecting a **pulse of energy** into a steady state turbulence field at $H=t=0$
And measure the velocity field at $H>0$.



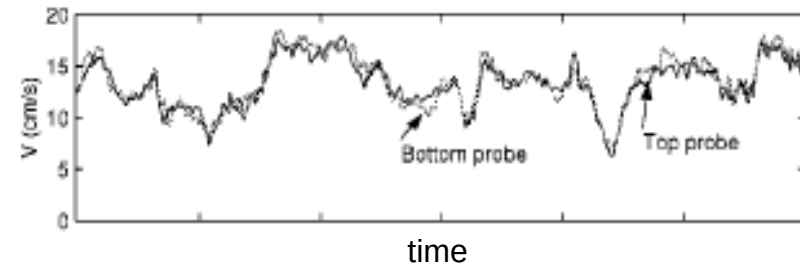
It is known:

Built up of large scales

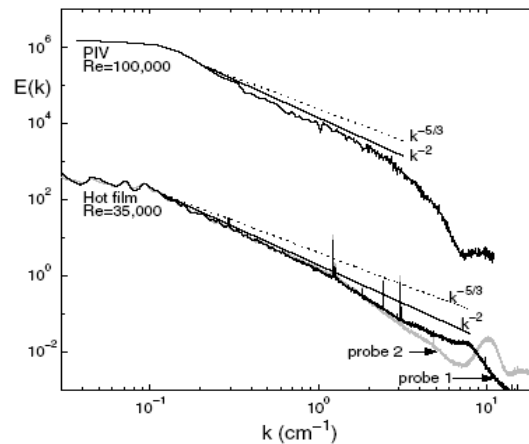
2D in the large scales (Baroud, Plapp and Swinney, 2003)

But also:

Different energy power spectra:



Baroud et. al. 2002



Smith and Waleffe 1999

