Stellar activity, accretion, and chronometry in the UV

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Rotating variables







Gyrochronology

- As stars age, they spin down their rotation from magnetic dragging
- In future, measuring rotational period will allow estimating ages of stars



Searching for variable stars with Kepler

- Full coverage of the period space
- 34,000 rotating variables
- Only 115 sq. deg field
 - Only sparce calibration data for gyrochronology





Searching for variable stars with TESS

- Almost complete sky coverage
- Only 27 day baseline
 - Difficult to measure rotation periods longer than 12 days
 - Missing 92% of variables

TESS 2-year sky coverage map





Searching for variable stars with ULTRASAT

- Best of both worlds:
 - 4 day cadence is perfect to measure periods of long period rotators
 - Area of coverage ~20 times of Kepler



Spot properties

- Only a handful of stars have independently derived spot temperatures
- Difference in color between ULTRASAT and TESS will significantly increase the census



Flares

- Flares are 15 times brighter in UV than in optical
- Evolution of flare properties as a function of age and mass





Eclipsing binaries

- Fundamental stellar laboratories
- Enable directly measuring mass, radius, Teff, etc
- Characterization of chromospheric activity in EBs is particularly valuable



Accretion in young stars

• ULTRASAT will provide all sky UV photometery



Accretion in young stars

- ULTRASAT will provide all sky UV photometery
- UV excess in young stars from accretion shocks





Accretion in young star:

- One of the most direct tracers of the accretion rate
- Require only photometry



Synergies with NASA UVEX mission **UVEX** Capabilities UVEX PSF: 2" **Small Magellanic** -68° WFC3/UVIS O 9V star Cloud GALEX SWIFT/UVOT **UVEX Spectrograph** 2° length 4.1 solar mass stripped star -70° GALEX PSF: 5.3" B 2V star -72° ULTRASAT PSF: 13" WNE-class Wolf-Rayet star **UVEX** Imager -74° 3.5° x 3.5° 18 GALEX 🗳 fraction of sky 5 arcmin UV Spectroscopy from 1150 to 2650 Å FUV 1^h30^m 2^h00^m 00^m 0^h30^m observed 20 NUV AB Magnitude GALEX 2003 - 2013 ULTRASAT 2025 → 8 × HST 1990 → Euclid 👄 deepei **ULTRASAT** YJH Swift 2004 → UVEX 2028 → **FUV** NUV UVEX Euclid 2023 → 28 UVEX: an all-sky ultraviolet Rubin 📛 Rubin 2024 → mission for the coming Roman 2025 → generation 30 104 2×10^{3} 5×10³ LIGO/Virgo/KAGRA 05 06 Wavelength (Å)

Summary

- All modes of operations of ULTRASAT are valuable for understanding of stellar evolution
- Short cadence
 - Flares
- Long cadence
 - Long period rotators
- All-sky imaging
 - Accretion of young stars
- UVEX synergy?